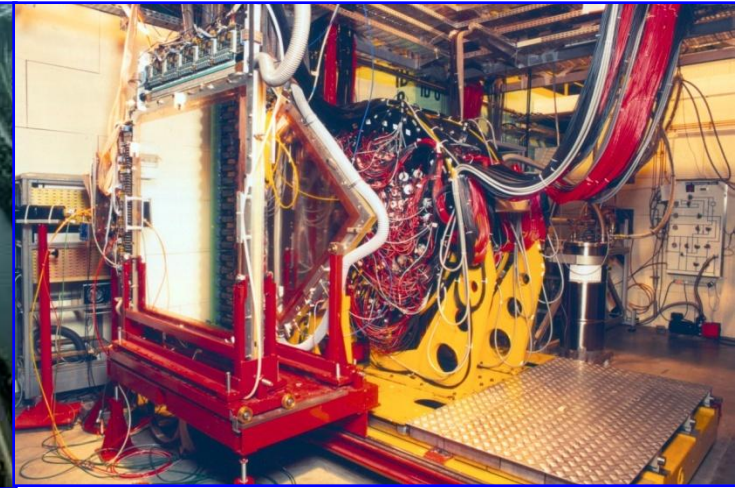
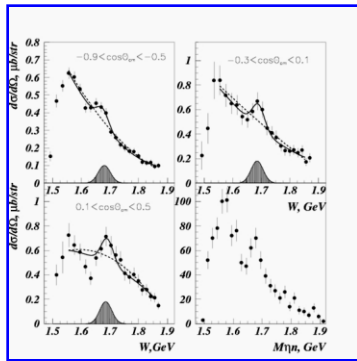


Recent Updates from GRAAL

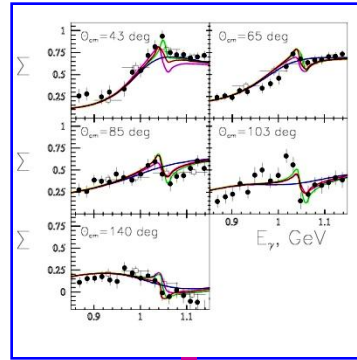


*Viacheslav Kuznetsov, PNPI,
in collaboration with Nuclear Physics Group
(University of Catania and INFN-Catania) and
Maxim Polyakov (Ruhr-Universitat-Bochum)*

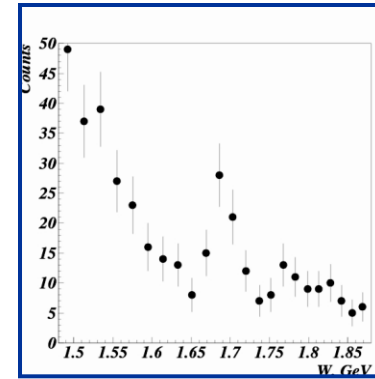
Graal $\gamma n \rightarrow \eta n$



Graal $\gamma p \rightarrow \eta p$

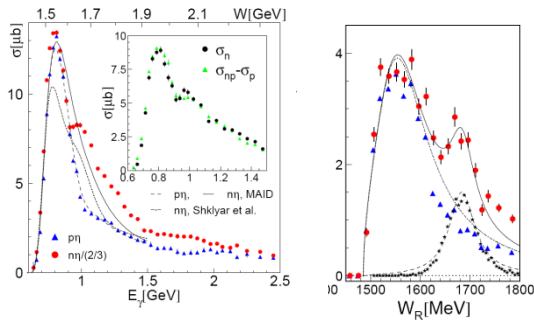
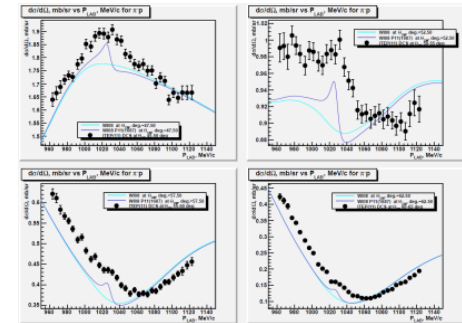


Graal $\gamma n \rightarrow \gamma n$

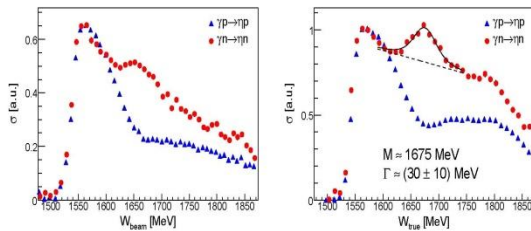


Neutron anomaly

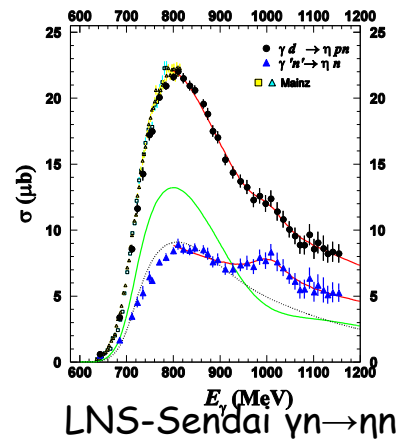
EPECUR $\pi^+ \rightarrow \pi^+$



CBELSA/TAPS $\gamma n \rightarrow \eta n$



Mainz $\gamma n \rightarrow \eta n$

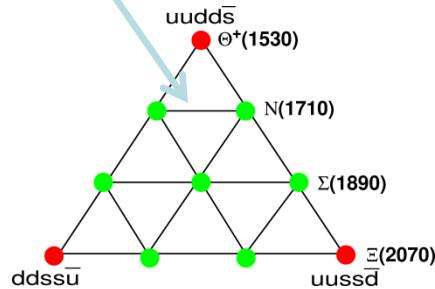


LNS-Sendai $\gamma n \rightarrow \eta n$

Neutron anomaly

Different interpretations

Narrow $N^*(1685)$
resonance
Several publications...



Interference of well-known wide isospin $\frac{1}{2}$ resonances ($S_{11}(1535)$, $S_{11}(1650)$, $P_{11}(1710)$...) Bonn-Gatchina Group (A.Sarantsev et al.,) Giessen Group (A. Mozel et al.,) and several others.
Widely discussed in literature!

Neutron
Anomaly

Citation: J. Beringer et al. (Particle Data Group). PR D86, 010001 (2012) (URL: <http://pdg.lbl.gov>)

$N(1685) \text{ ? ?}$

$I(J^P) = \frac{1}{2}(??)$ Status: *

OMITTED FROM SUMMARY TABLE

There is a small literature (which we do not try to cover) on this possible narrow state. See KUZNETSOV 11A, MART 11, and the other papers for further references. This state does not gain status by being a sought-after member of a baryon anti-decuplet.

N(1685) MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
~ 1670	JAEGLE 11	CBTP	$\gamma d \rightarrow \eta n (p)$
~ 1685	KUZNETSOV 11	GRAL	$\gamma d \rightarrow \gamma n (p)$
~ 1680	KUZNETSOV 07	GRAL	$\gamma d \rightarrow \eta n (p)$

N(1685) WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
~ 25	JAEGLE 11	CBTP	$\gamma d \rightarrow \eta n (p)$
<< 30	KUZNETSOV 11	GRAL	$\gamma d \rightarrow \gamma n (p)$
< 30	KUZNETSOV 07	GRAL	$\gamma d \rightarrow \eta n (p)$

••• We do not use the following data for averages, fits, limits, etc. •••

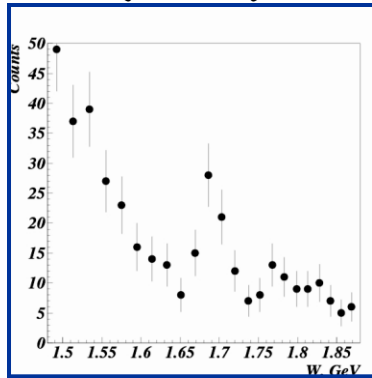
N(1685) REFERENCES

JAEGLE 11	EPJ A47 40	I. Jaegle et al.	(CBELSA/TAPS Collab.)
Also	PR 109 252002	I. Jaegle et al.	(CBELSA/TAPS Collab.)
KUZNETSOV 11	PR C83 022201	V. Kuznetsov et al.	(GRAAL Collab.)
KUZNETSOV 11A	JETPL 94 503	V. Kuznetsov, M.V. Polyakov, M. Thurnann	(JHRM+)
MART 11	PR D83 040015	T. Mart	(U. Indonesia)
KUZNETSOV 07	PL B1647 23	V. Kuznetsov et al.	(GRAAL Collab.)

- Intermediate sub-threshold $K\Sigma$ and KA states (cusp effect)
M.Doring, K. Nakayama, PLB683, 145 (2010),
nucl-th/0909.3538

Comments on the interpretation of neutron anomaly in terms of the specific interference of well-known resonances.

V.Kuznetsov et al., Phys.Rev. C83 (2011) 022201



One major challenge for this interpretation is the observation of a narrow enhancement at $W \sim 1.68$ GeV in Compton scattering on the neutron ($\gamma n \rightarrow \gamma n$),

Comments on “Interference phenomena in the $J^P = 1/2^-$ - wave in η photoproduction” by A.V. Anisovich, E. Klempt, B. Krusche, V.A. Nikonov, A.V. Sarantsev, U. Thoma, D. Werthmuller, [arXiv:1501.02093v1](https://arxiv.org/abs/1501.02093v1) [nucl-ex].

Viacheslav Kuznetsov

¹Petersburg Nuclear Physics Institute, Gatchina, 188300, St. Petersburg, Russia

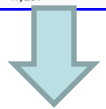
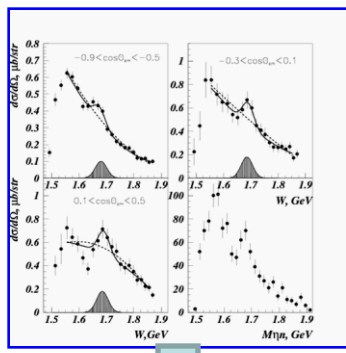
The authors of Ref. [1] claimed that “... narrow structure observed in the excitation function of $\gamma n \rightarrow \eta n$ can be reproduced fully with a particular interference pattern in the $J^P = 1/2^-$ partial wave...” while a narrow structure in Compton scattering off the neutron is “...a stand-alone observation unrelated to the structure observed in $\gamma n \rightarrow \eta n$...”. The source for the second statement may be a simple numerical error. If so, the interpretation of the narrow structure in $\gamma n \rightarrow \eta n$ as interference effects in the $J^P = 1/2^-$ -wave and some conclusions from Ref. [1] are questionable.

In accordance with (corrected) calculations by A. Anisovich et al., the total cross-section of $N^(1685)$ should be 10-25 nb. Therefore, if $N^*(1685)$ does exist, its peak should be clearly seen in the Compton cross-section on the neutron.*

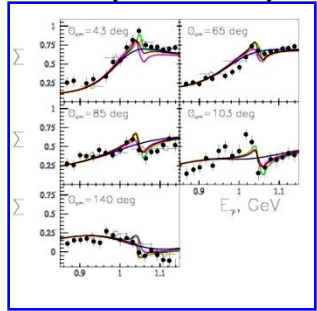
The observation of the peak in Compton scattering on the neutron is in fact refutes the explanation of the neutrons anomaly in terms of interference phenomena. This interference cannot generate a peak in eta photoproduction, which is governed by isospin-1/2 resonances, simultaneously generate the same peak in Compton scattering, which is governed by isospin-1/2 and isospin-3/2 resonances, and generate neither of peak in pion photoproduction on the neutron, which is governed by the same resonances as Compton scattering.

One benchmark signature of $N^(1685)$ is the strong photoexcitation on the neutron and suppressed (but not zero) photoexcitation on the proton. Such a resonance would appear in cross section on the proton as a minor peak/dip structure which may not be resolved in experiment. On the contrary, its signal may be amplified in polarization observables being amplified due to the interference with other resonances.*

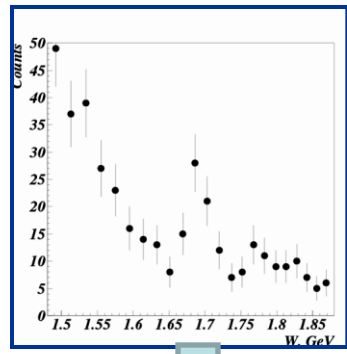
Cross section $\gamma n \rightarrow \eta n$



Beam asymmetry $\gamma p \rightarrow \eta p$



Cross section $\gamma n \rightarrow \gamma n$



Beam asymmetry $\gamma p \rightarrow \gamma p$



Compton scattering $\gamma p \rightarrow \gamma p$ on the proton at GRAAL: Surprising results

Evidence for narrow resonant structures at $W \approx 1.68$ GeV and $W \approx 1.72$ GeV in real Compton scattering off the proton

V. Kuznetsov,^{1,*} F. Mammoliti,^{2,3} V. Bellini,^{2,3} G. Gervino,^{4,5} F. Ghio,^{6,7} G. Giardina,⁸ W. Kim,⁹ G. Mandaglio,^{2,8}
M. L. Sperduto,^{2,3} and C. M. Sutura^{2,3}

¹*Petersburg Nuclear Physics Institute, 188300, Gatchina, Russia*

²*INFN - Sezione di Catania, via Santa Sofia 64, I-95123 Catania, Italy*

³*Dipartimento di Fisica ed Astronomia, Università di Catania, I-95123 Catania, Italy*

⁴*Dipartimento di Fisica Sperimentale, Università di Torino, via P.Giuria, I-00125 Torino, Italy*

⁵*INFN - Sezione di Torino, I-10125 Torino, Italy*

⁶*INFN - Sezione di Roma, piazzale Aldo Moro 2, I-00185 Roma, Italy*

⁷*Istituto Superiore di Sanità, viale Regina Elena 299, I-00161 Roma, Italy*

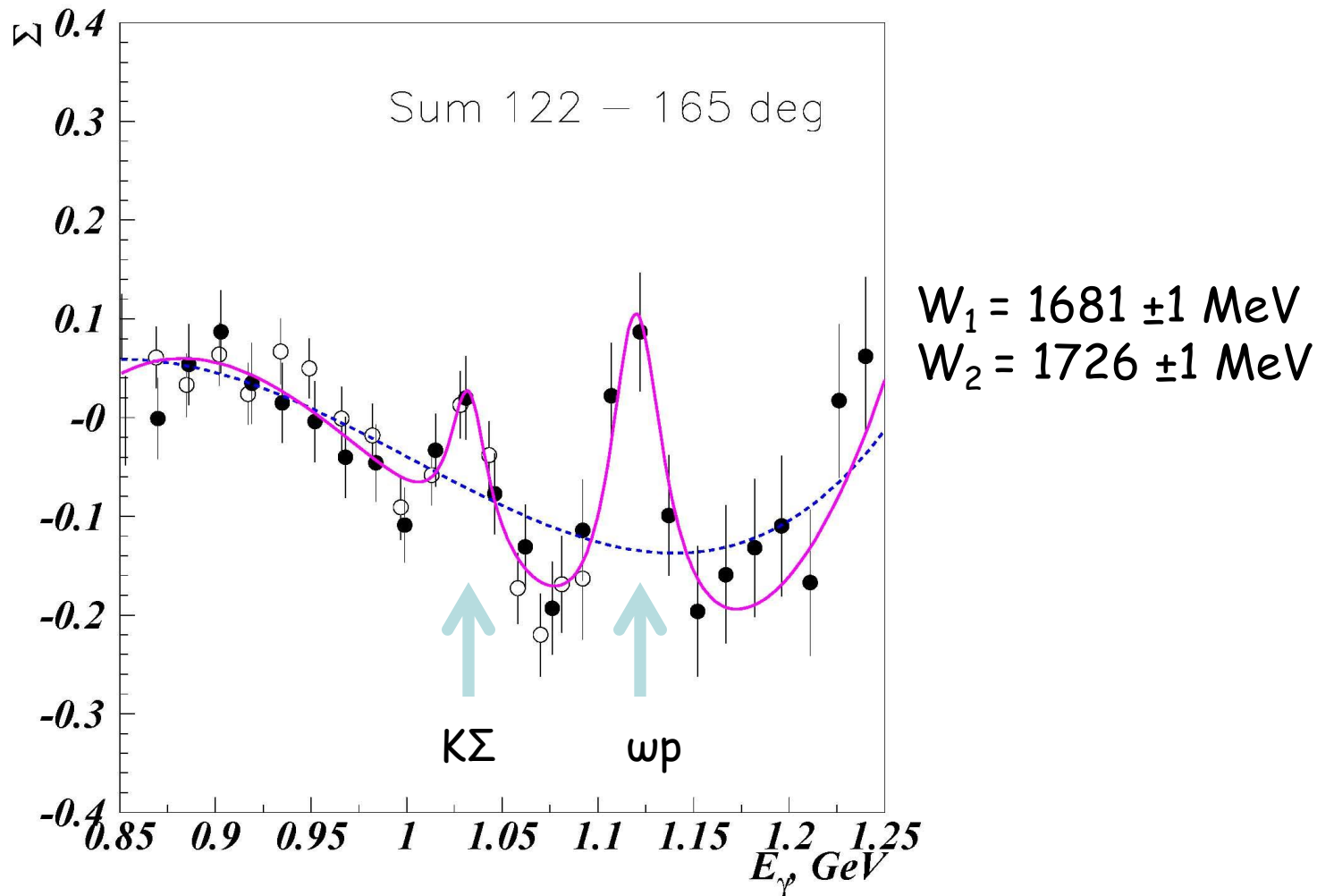
⁸*Dipartimento di Fisica e di Scienze della Terra - Università di Messina, salita Sperone 31, 98166 Messina, Italy*

⁹*Kyungpook National University, 702-701, Daegu, Republic of Korea*

(Received 18 January 2015; published 6 April 2015)

First measurement of the beam asymmetry Σ for Compton scattering off the proton in the energy range $E_\gamma = 0.85$ – 1.25 GeV is presented. The data reveal two narrow structures at $E_\gamma = 1.036$ and $E_\gamma = 1.119$ GeV. They may signal narrow resonances with masses near 1.68 and 1.72 GeV, or they may be generated by the sub-threshold $K\Lambda$ and ωp production. Their decisive identification requires additional theoretical and experimental

Two narrow ($\Gamma \sim 20$ MeV) structures at $W \sim 1.68$ and $W \sim 1.72$ GeV in the beam asymmetry data for Compton scattering off the proton at GRAAL



Comment on “Evidence for narrow resonant structures at $W \approx 1.68$ GeV and $W \approx 1.72$ GeV in real Compton scattering off the proton”

D. Werthmüller,^{1,2} L. Witthauer,² D. I. Glazier,¹ and B. Krusche²

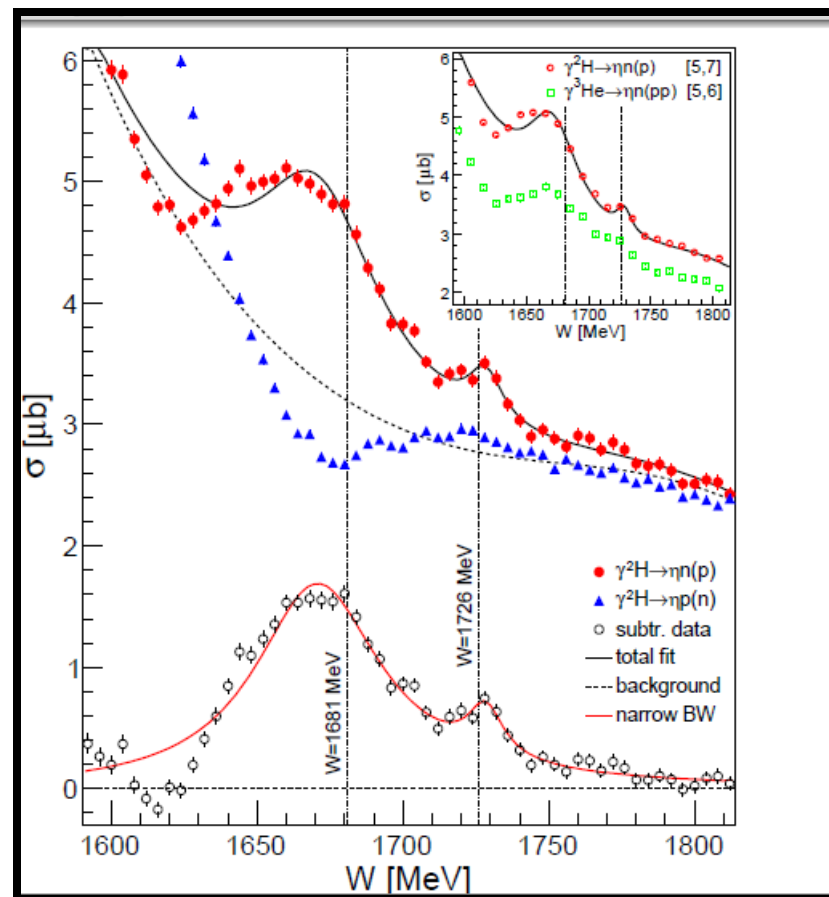
¹*School of Physics and Astronomy, University of Glasgow, Glasgow G12 8QQ, Scotland, United Kingdom*

²*Departement Physik, Universität Basel, CH-4056 Basel, Switzerland*

(Received 8 July 2015; published 11 December 2015)

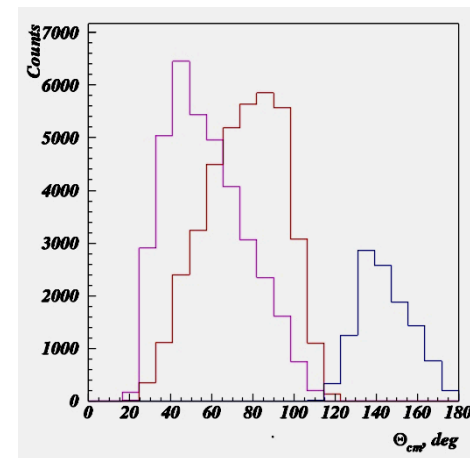
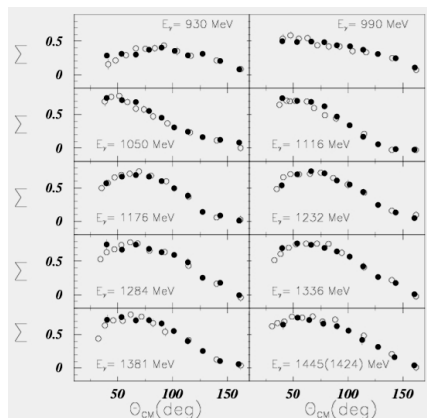
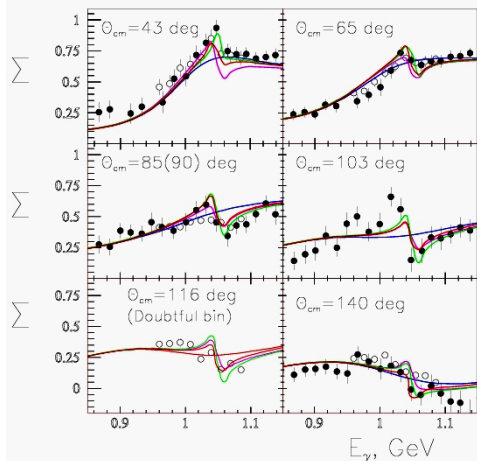
We comment on the statement by Kuznetsov *et al.* that the structure around $W = 1.72$ GeV seen in the beam asymmetry in Compton scattering off the proton is not observed in the total cross section of η photoproduction on the neutron.

Observation of two narrow structure at $W \sim 1.68$ and $W \sim 1.72$ GeV in $\gamma n \rightarrow \eta n$ at A2@MaMiC and CBELSA/TAPS



Revised analysis of the beam asymmetry Σ for eta photoproduction off the proton $\gamma p \rightarrow \eta p$

Previous GRAAL data



*V. Kuznetsov, M.V.P., et al.,
Acta Physica Polonica , 39
(2008) 1949*

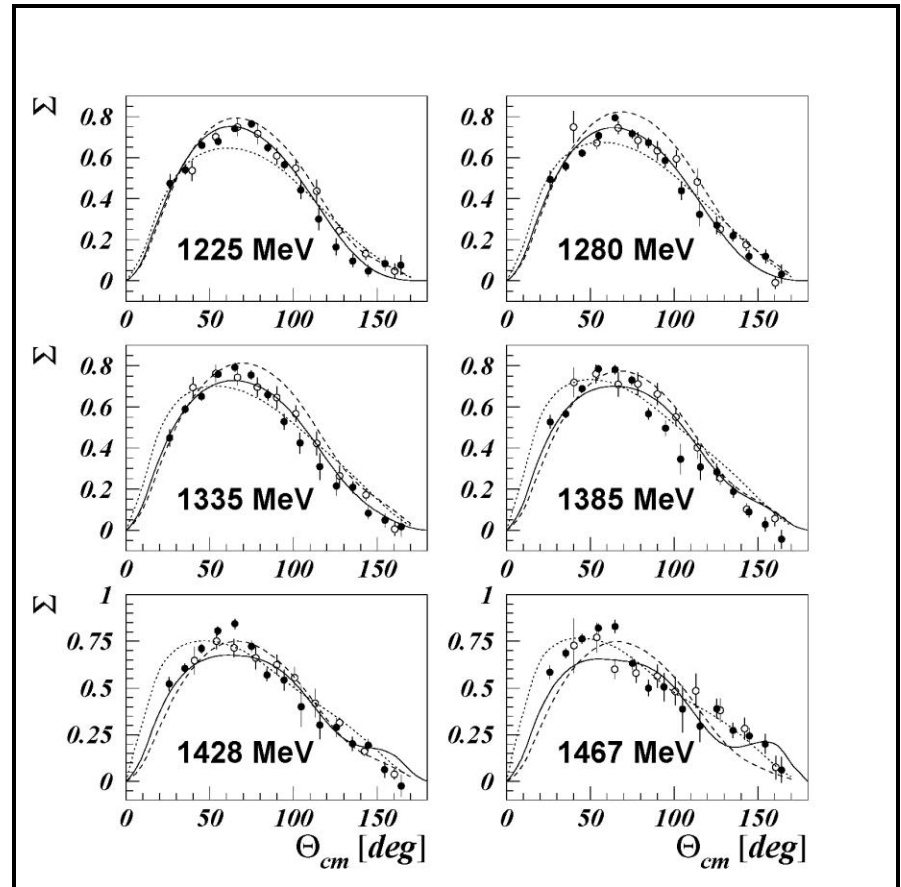
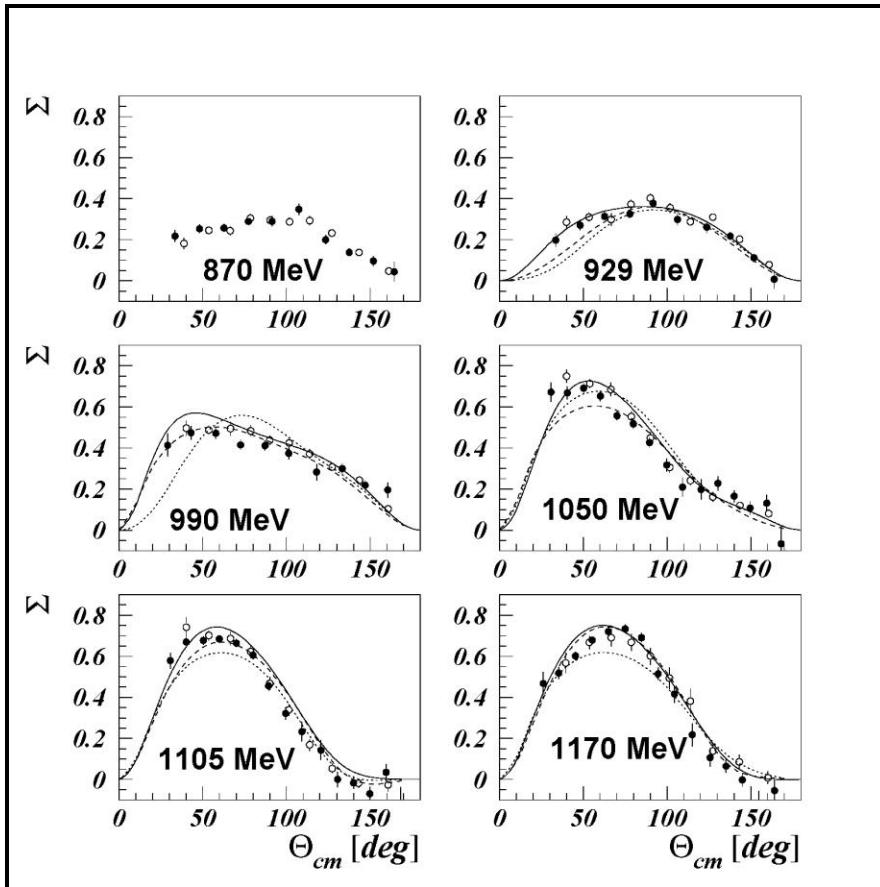
*V. Kuznetsov, M.V.P., JETP
Lett., 88 (2008) 347*

*Fine energy binning
Observation of resonant
structure at $W \sim 1.68$ GeV,
 $\sim 10\%$ of available statistics
was used*

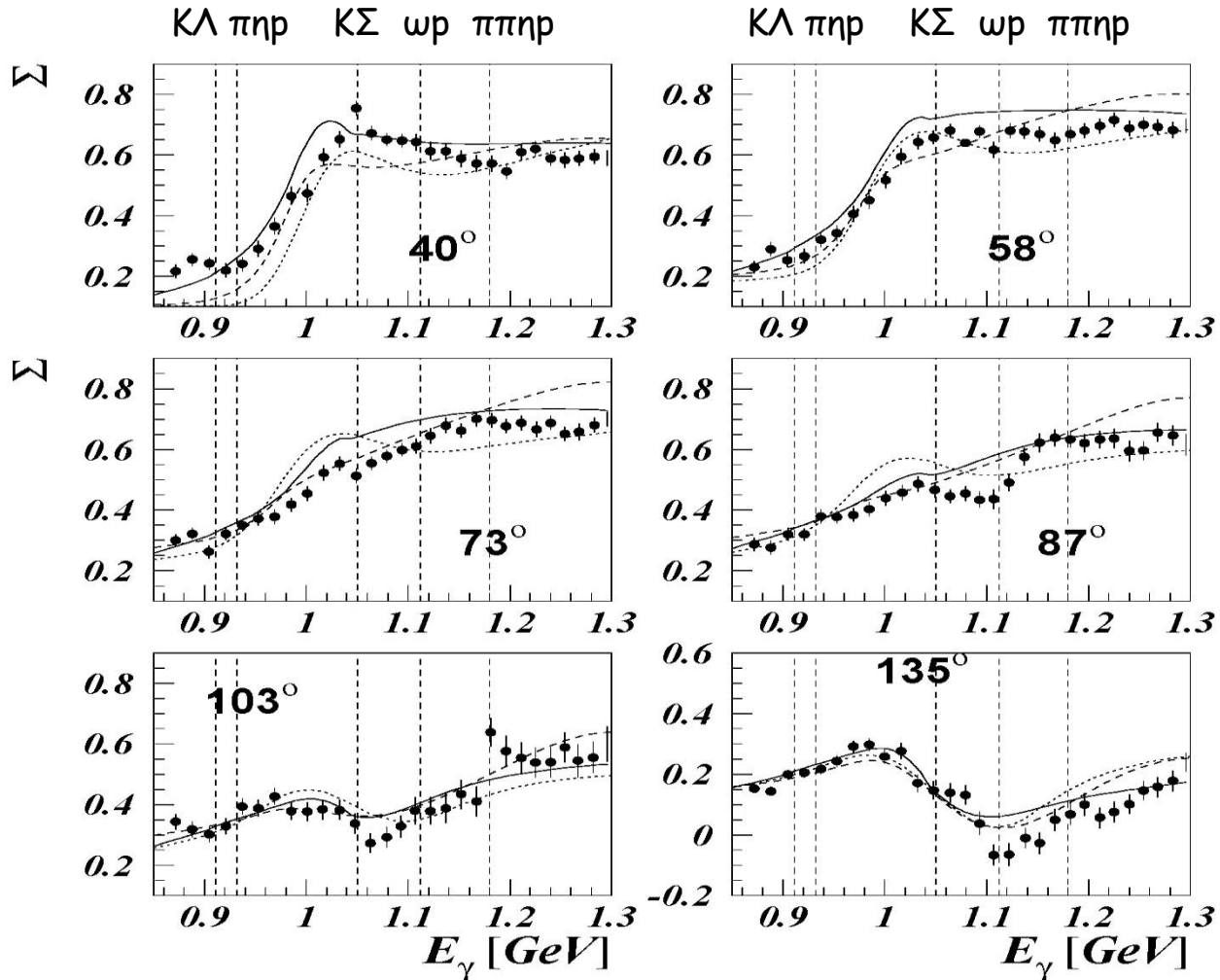
*O. Baratlini et al., Eur.Phys.J.
A33 (2007) 169-184
(black points)
Wide energy binning/
Bugs in analysis. The critics
remains unreplied during 8
years.*

Wide energy binning

New data - black circles, old data
(O.Baratlini et al., Eur.Phys.J. A33 (2007) 169-184) - open circles
Solid line - Boga, dashed line - SAID, dotted line -Eta-Maid.



New data with fine energy binning



Summary:

- *Several narrow peculiarities close to meson-nucleon thresholds are observed;*
- *This favors the assumption of the cusp effect. However, there are open questions: i) why some of them are better on the neutron; ii) Could cusps occur in Compton scattering and not manifest themselves; in pion photoproduction;*
- *There might be other explanations (for example, formation of bound meson-nucleon states)?*

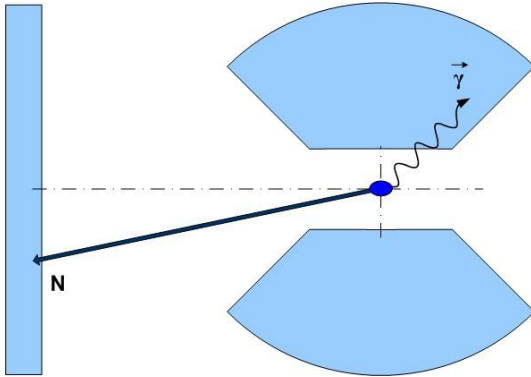
Critical issues:

- *To provide more experimental data;*
- *To perform a combined PWA of new data. Cusps are a priori S-wave effects.....*

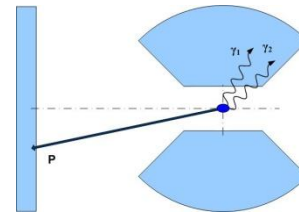
Thank you for your attention!

Compton scattering and the π^0 background.

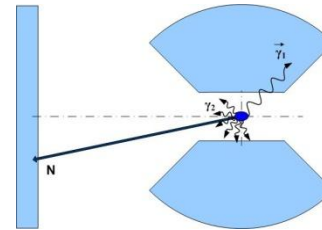
Compton scattering



π^0 background



Symmetric π^0 decay, two photon hits are mixed in one BGO cluster.

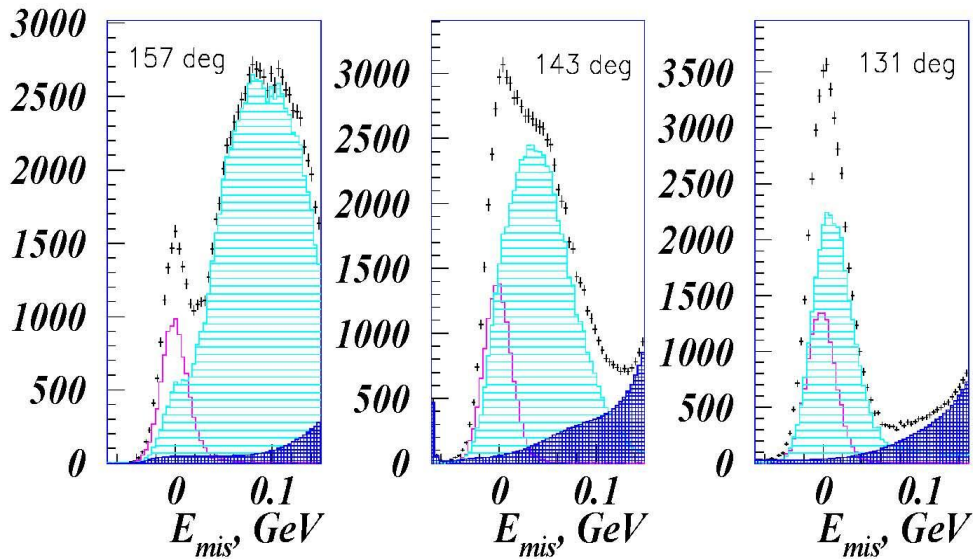


Asymmetric π^0 decays, the second low-energy photon may not be detected.

Main identification parameter is
Missing energy

$$E_{mis} = E_{\gamma} - E_{\gamma'} - T_p(\theta_p)$$

Simulation and Data

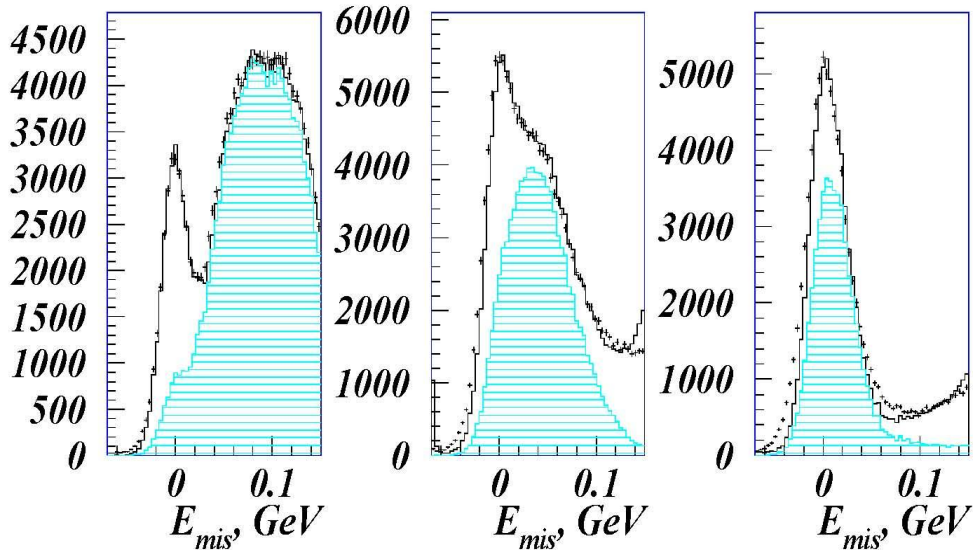


Simulations:

Magenta: Compton

Green : the pion background

Blue: other reactions



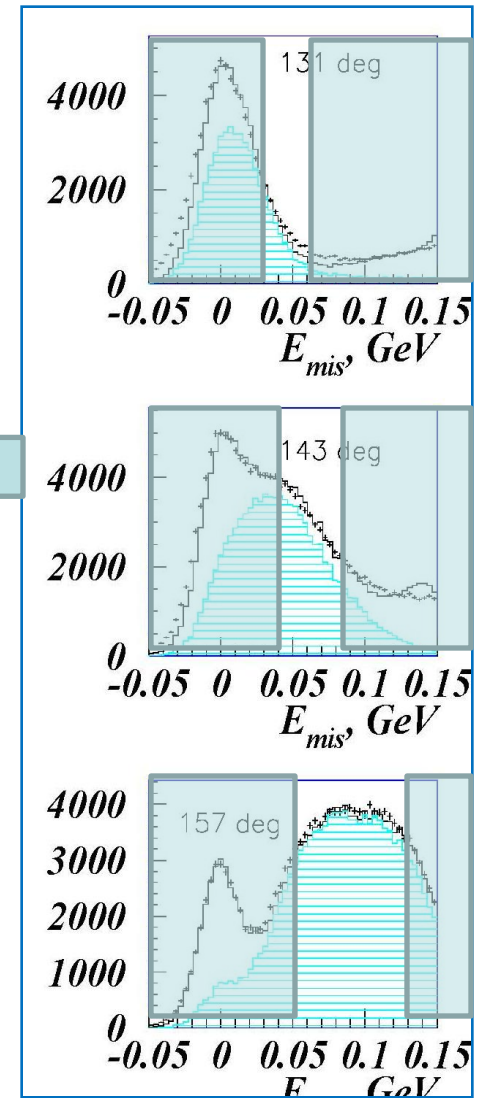
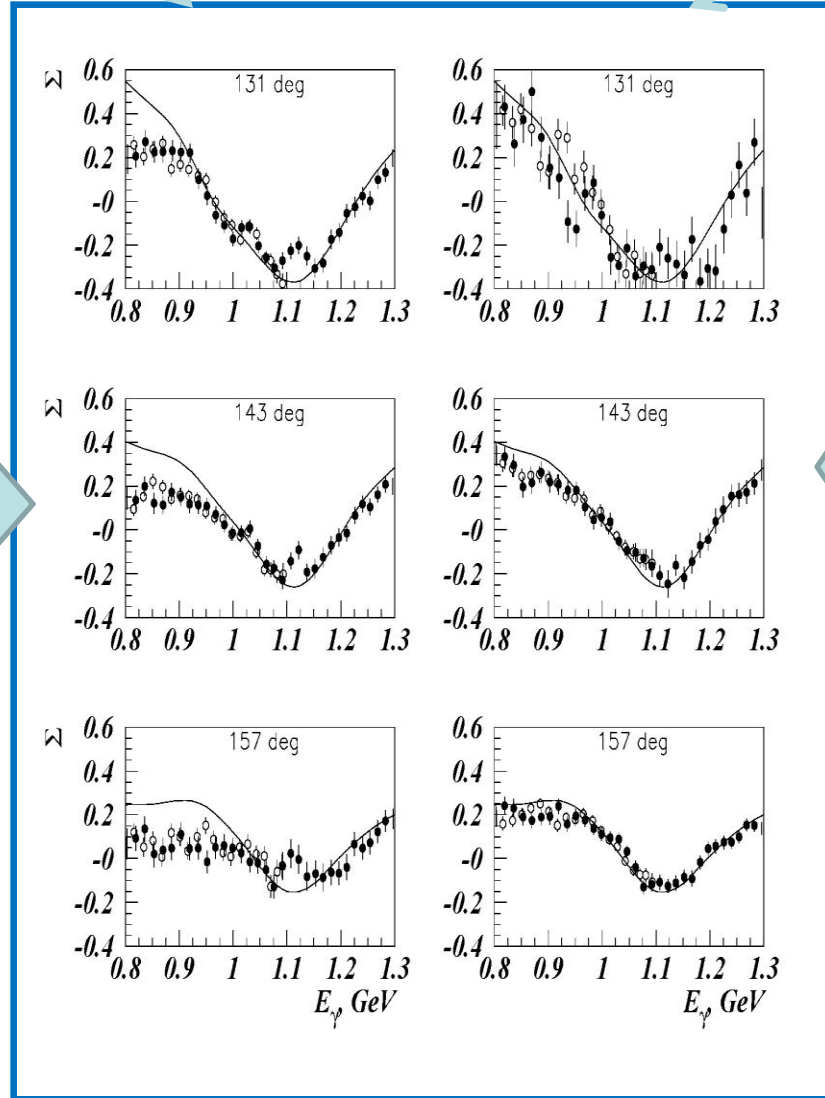
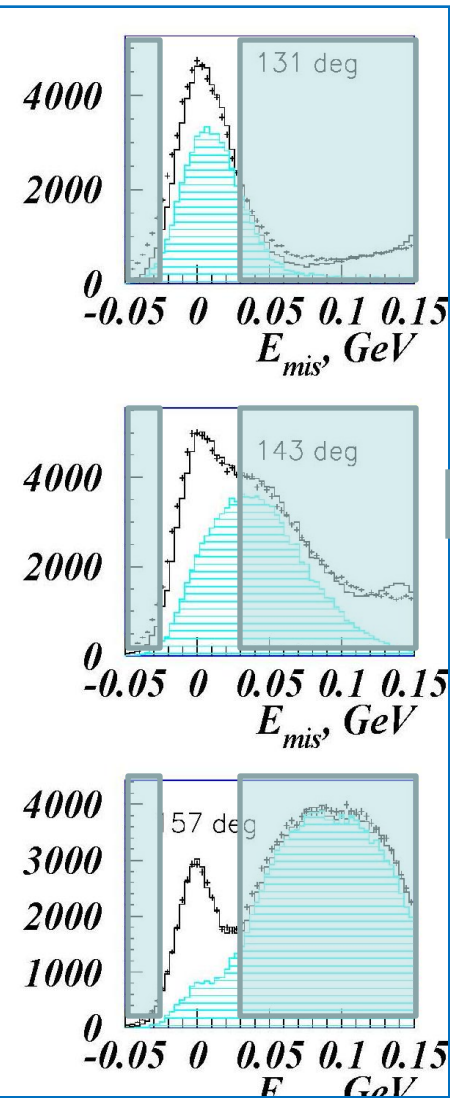
Data:

Green is the estimated contamination of pion photoproduction.

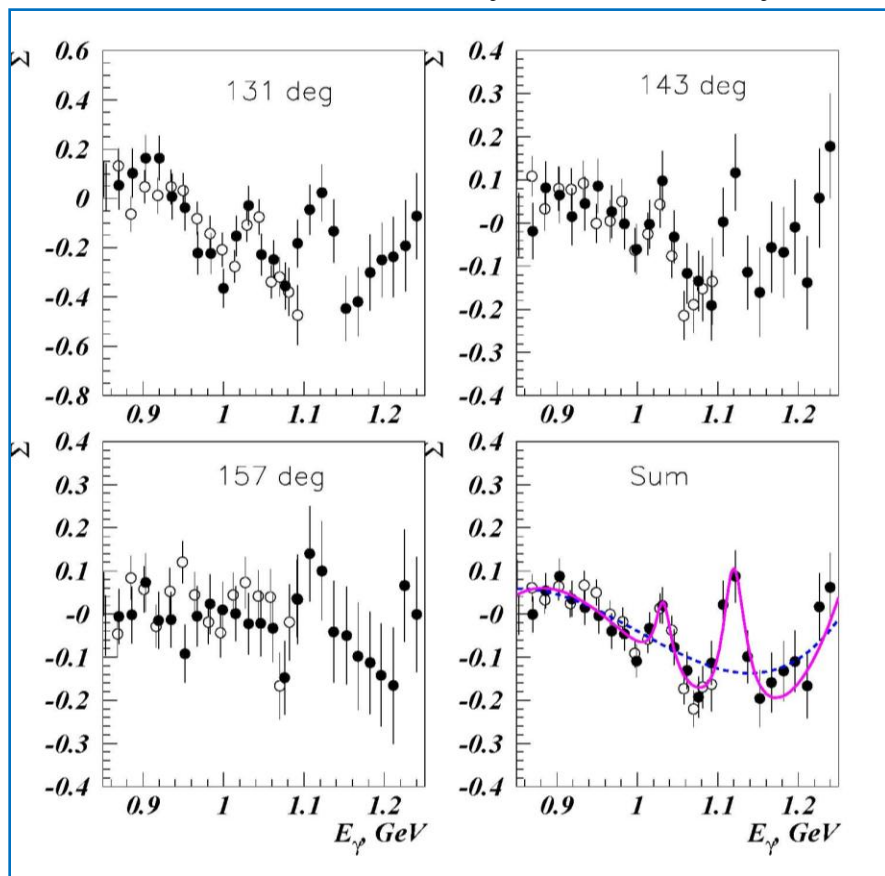
Beam asymmetry with the main and side-band cuts.

Mixture of Comptons and pions

Mostly pions



Beam Asymmetry for Compton scattering



Two narrow resonant structures!

Fits:

*4-order polynomial
(background hypothesis)*

$$\chi^2 = 77/39$$

4-order polynomial + BW1 + BW2

$$BW = \frac{(E_\gamma - E_r) \sin(\varphi) + \Gamma \cos(\varphi)}{(E_\gamma - E_r)^2 + \frac{\Gamma^2}{4}}$$

$$\chi^2 = 27/33$$

Results

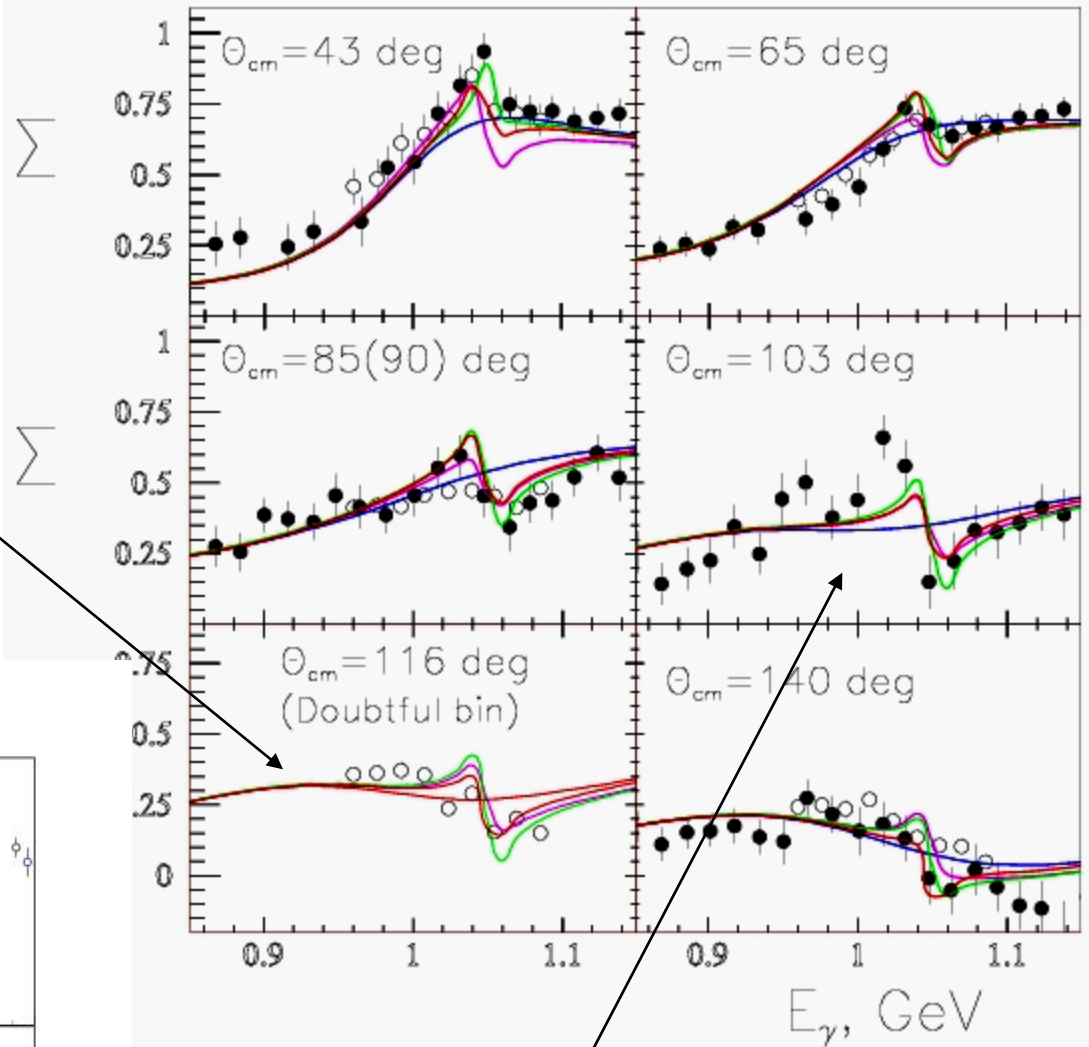
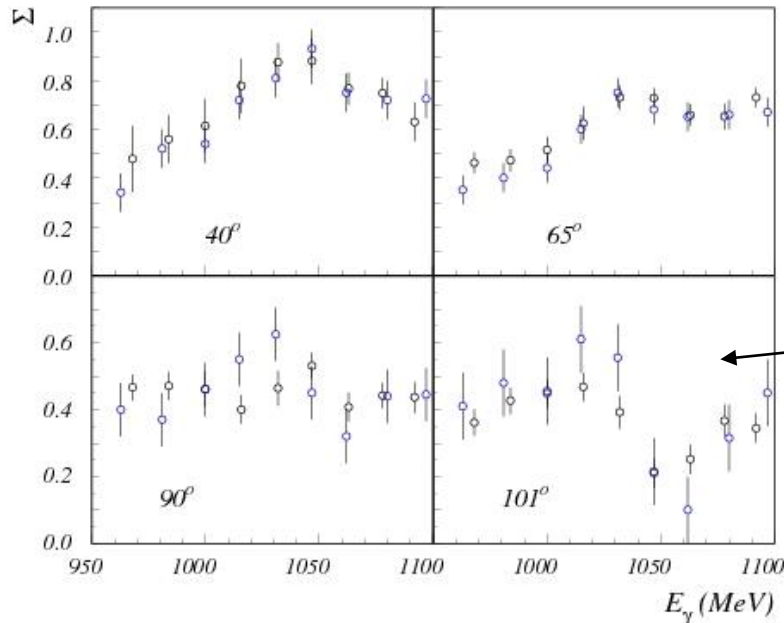
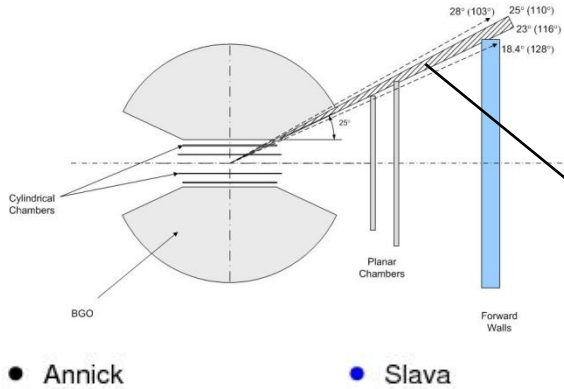
$E_1 = 1.036 \text{ GeV}$	$\Gamma_1 = 25 \text{ MeV}$	$W_1 = 1.681 \pm 0.006 \text{ GeV}$	$\Gamma_{w1} = 16 \pm 6 \text{ MeV}$
$E_2 = 1.119 \text{ GeV}$	$\Gamma_2 = 35 \text{ MeV}$	$W_2 = 1.726 \pm 0.007 \text{ GeV}$	$\Gamma_{w1} = 21 \pm 8 \text{ MeV}$

Summary:

- Two resonant structures at $W \sim 1.68$ and $W \sim 1.72$ GeV are observed;
- One more challenge for the explanation of the "neutron anomaly" in terms of the interference of well-known resonances;
- $W \sim 1.68$ GeV and $W \sim 1.72$ GeV are close to the thresholds of $K\Lambda$ and ωp photoproductions. This favors the assumption of the cusp effect. However, there are open questions: i) why it is not seen in pion photoproduction; ii) Could it occur in Compton scattering; iii) Why the structure at $W \sim 1.72$ GeV is not seen in η photoproduction?
- The only explanation that accommodates all experimental findings at $W \sim 1.68$ GeV is the existence of $N^*(1685)$ resonance. The observation of the second structure at $W \sim 1.72$ GeV adds more puzzle.....

Thank you for your attention!

Comparison of O.Bartalini et al. (open circles) and our results (black circles). Main difference is at 103/116 deg.



The same structure at 103 deg!

Comparison of preliminary results done by A.Lleres (A.Lleres, private communication (E-mail from Feb 5, 2007).

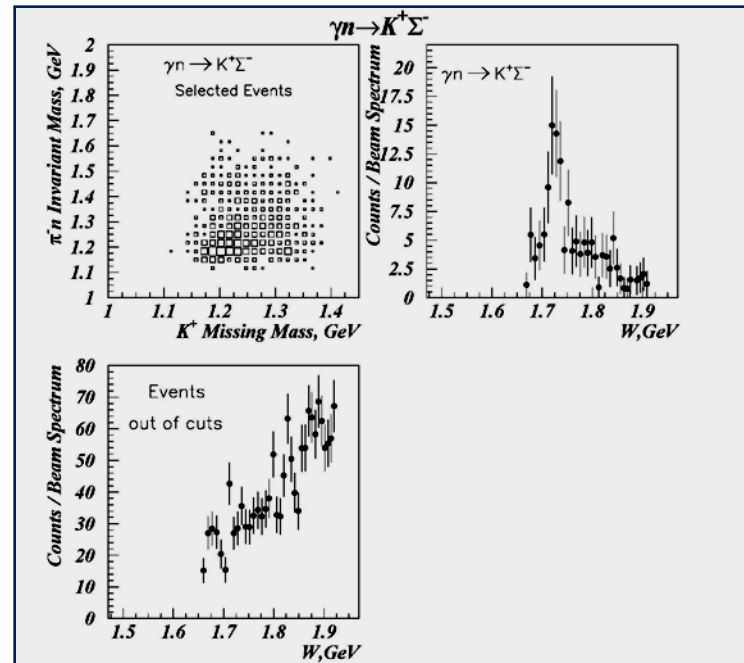
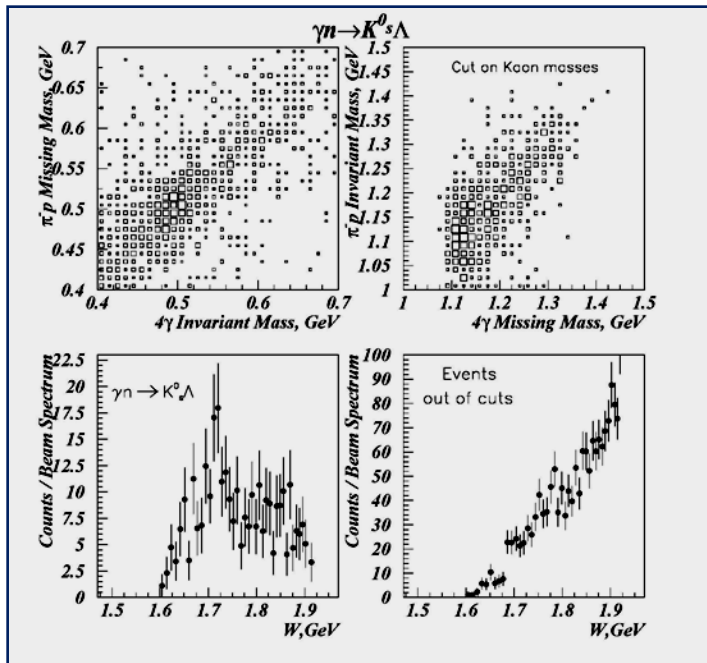
et al., NSTAR2007, September 2007

Resonant Structure at $W \sim 1.72$ GeV?

Some other results



Old (2003) preliminary GRAAL Analysis (remains uncompleted)
 V.Kuznetsov for the GRAAL Collaboration, Talk at Worksho "Pentaquarks
 2004", Trento, February 2004.

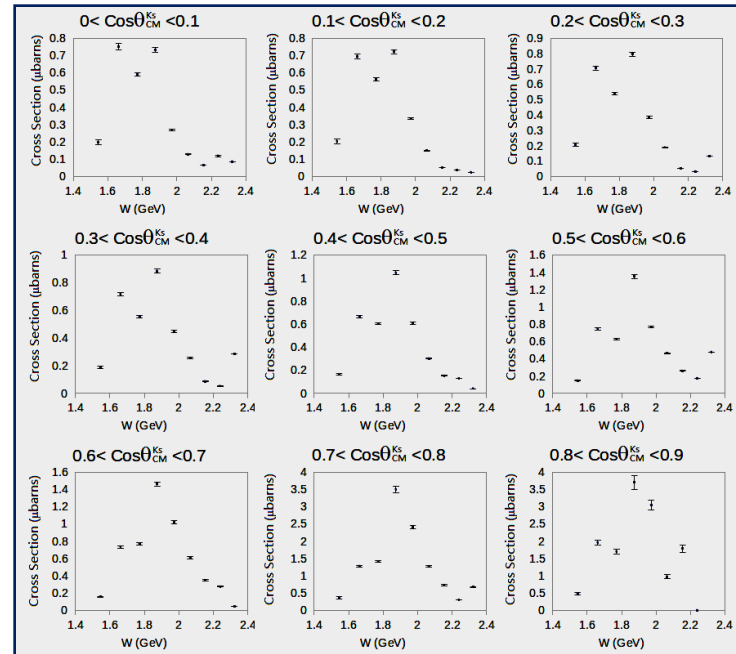
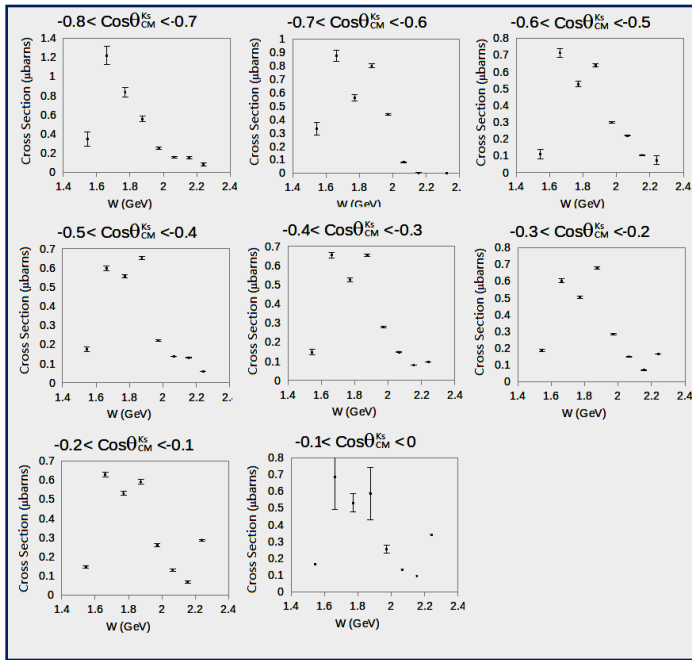


Peak near $W \sim 1.72 \text{ GeV}$?

Preliminary data on $\gamma n \rightarrow K^0_s \Lambda$ from CLAS

Talk of Taylor at NSATR2013 Workshop

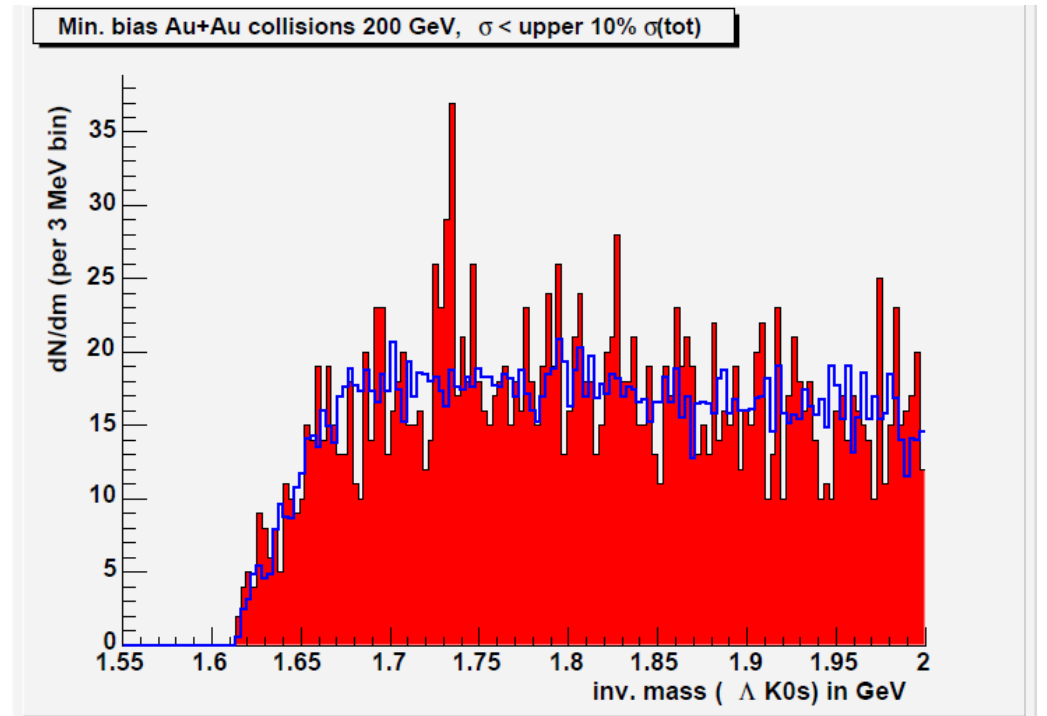
Peak at 1.7 GeV!



Old (2003) preliminary results from STAR

S. Kabana for the STAR Collaboration, Talk at Worksho "Pentaquarks 2004", Trento, February 2004.

S. Kabana for the STAR Collaboration,
PoS of 20th Winter Workshop on Nuclear Dynamics Trelawny
Beach, Jamaica March 15{20, 2004)



SAID PWA

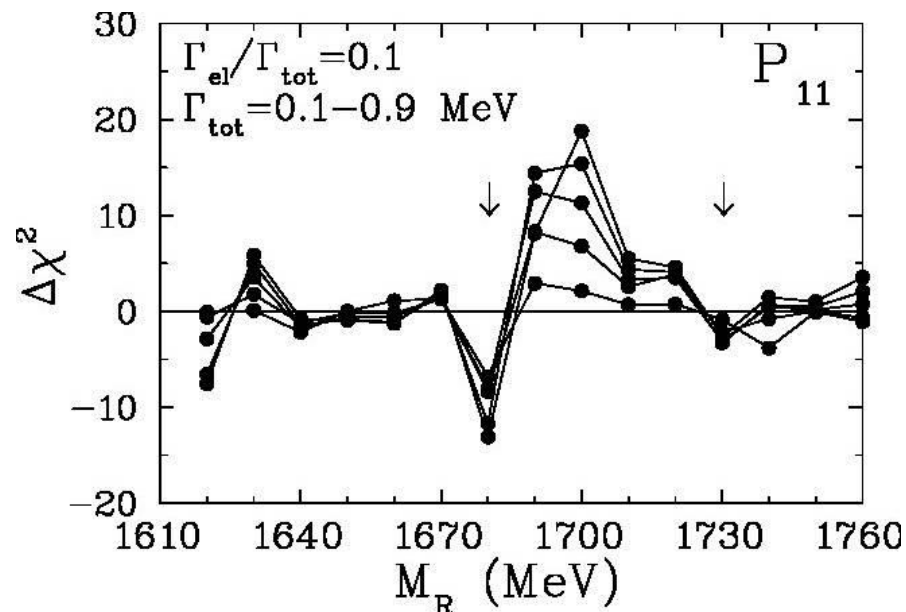
R.Arndt, Ya.Azimov, M.Polyakov, I.Strakovsky, R.Workman

“Nonstrange and other flavor partners of the exotic θ^+ baryon”

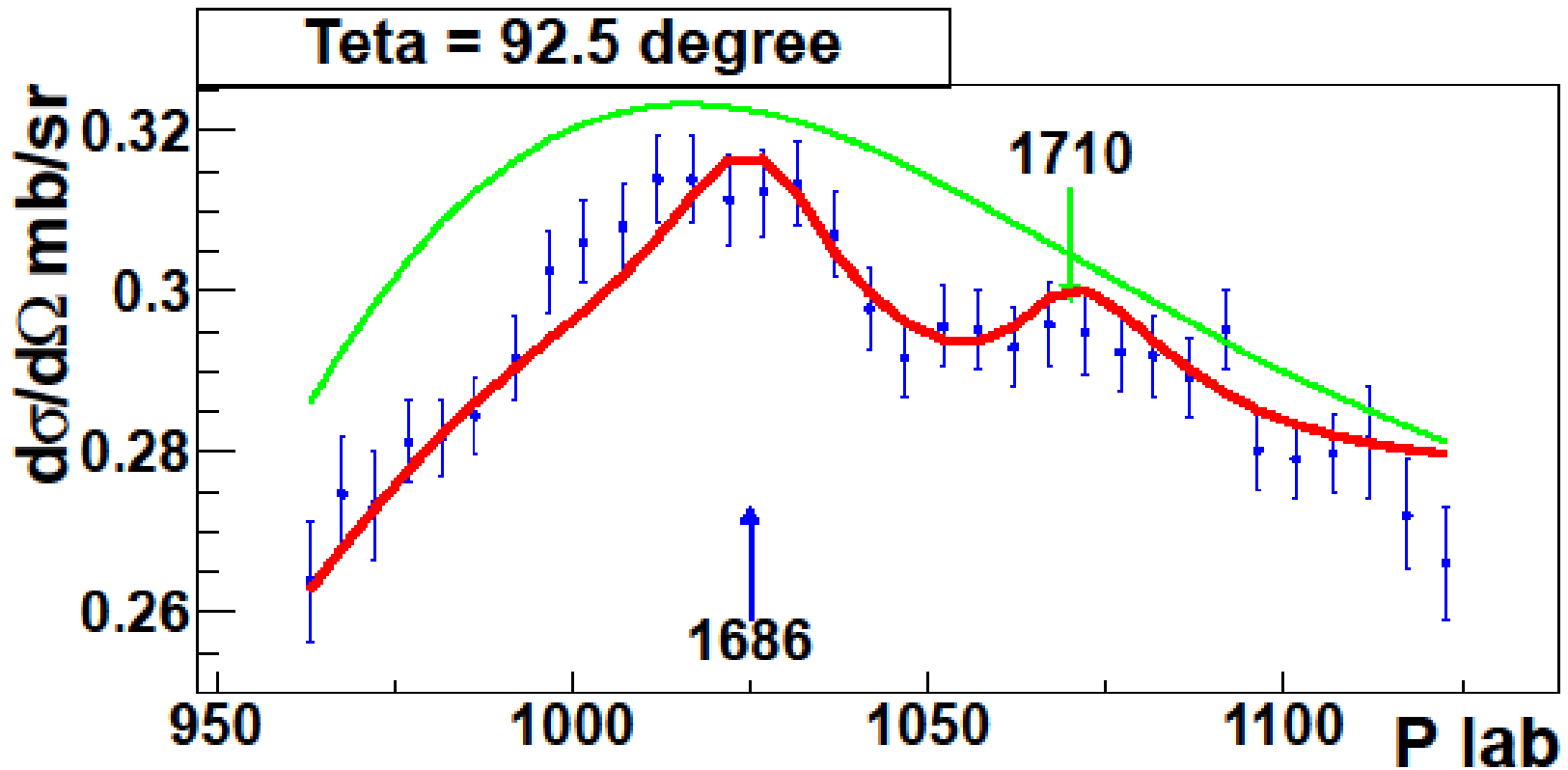
Phys.Rev. C69 (2004) 035208

Nucl-th/0312126;

“... given our present knowledge of the θ^+ , the state commonly known as the N(1710) is not the appropriate candidate to be a member of the antidecuplet. Instead we suggest candidates with nearby masses, N(1680) (more promising) and/or N(1730) (less promising, but not excluded). Our analysis suggests that the appropriate state should be rather narrow and very inelastic...”



Recent updates from EPECUR $\pi^-p \rightarrow \pi^-p$
A. Gridnev, Private Communication



Interpretations:

Narrow resonance

- Y. Azimov, V. Kuznetsov, M. Polaykov, and I. Strakovsky, Eur. Phys. J. A **25**, 325, 2005.
- A. Fix, L. Tiator, and M. Polyakov, Eur. Phys. J. A **32**, 311, 2007.
- K.S. Choi, S.I. Nam, A. Hosaka, and H-C. Kim, Phys. Lett. B **636**, 253, 2006.
- K.S. Choi, S.I. Nam, A. Hosaka, and H-C. Kim, Prog. Theor. Phys. Suppl. **168**, 97, 2008.
- G.S. Yang, H.S. Kim, Arxiv:1204.5644

• Interference of Known resonances V. Shklyar, H. Lenske, U. Mosel, PLB650 (2007) 172 (Giessen group); A. Anisovich et al. EPJA 41, 13 (2009), hep-ph/0809.3340 (Bonn-Gatchina group); X.-H. Zong and Q. Zhao, Arxiv:1106.2892

• Intermediate sub-threshold meson-nucleon state

M. Doring, K. Nakayama, PLB683, 145 (2010), nucl-th/0909.3538

Thank you for your attention!